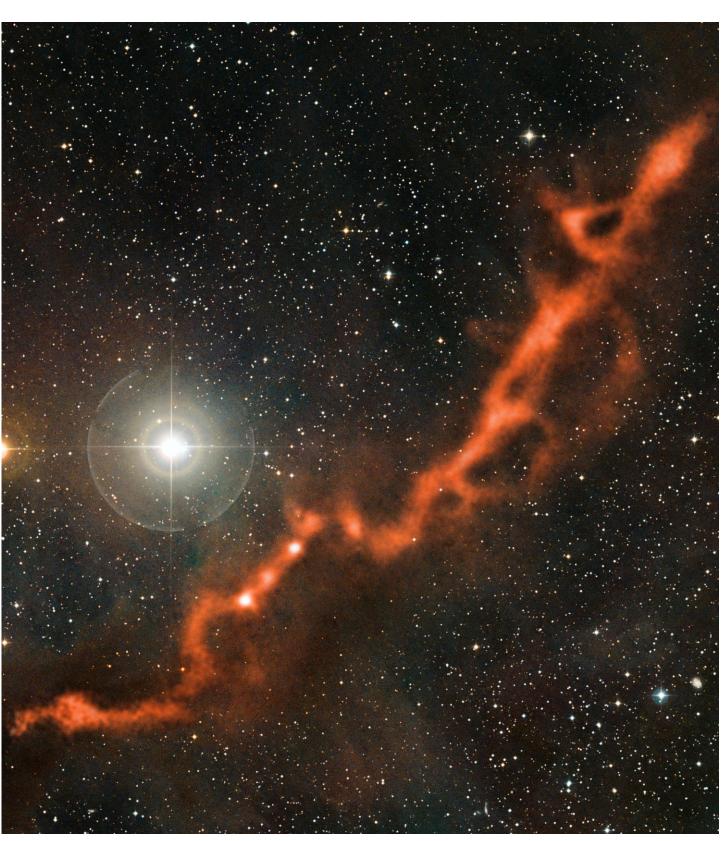
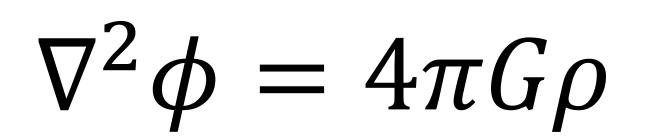


Motivation

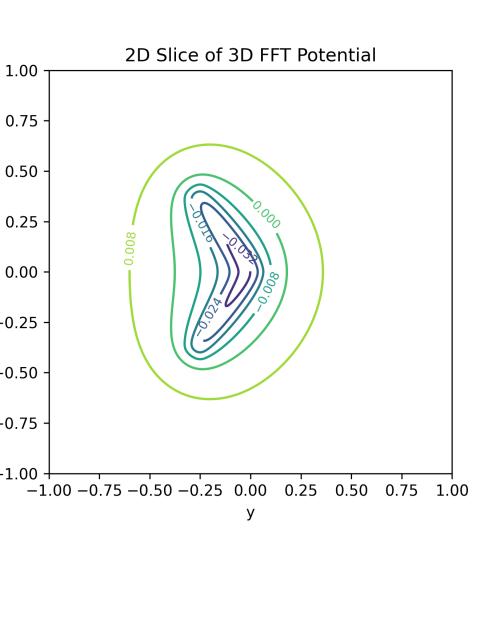
- Filaments are one structure of molecular gas where star formation is observed.
- The collapse modes and evolution of finite bent filaments is not yet completely understood, and the current models do not completely align with observed star formation.
- There have been models of straight finite filaments where the expected conditions occur, a pile-up of matter at the ends; however, star formation is not always seen in the ends.
- \succ The structure of a straight filament supports more formation at the ends, so there must be other factors, such as a bend or kink or smoothing of the ends of the filament. > My goal: model bent
- filaments and see if there is a point where the bend dominates



Filament Modeling



Solved Poisson's equation for gravity numerically using Fast Fourier Transforms to find the gravitational potentials 0.50 - $\triangleright \phi$ is the scalar potential 0.25 -N 0.00 · $\succ G$ is the gravitational constant -0.25 - $\triangleright \rho$ is the initial surface density -0.50 - \blacktriangleright Accelerations can than be found by -0.75taking the negative gradient of this result



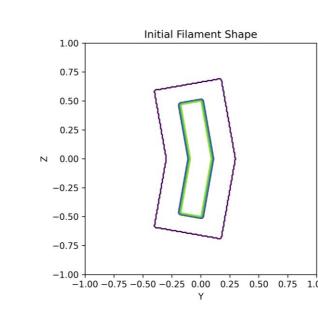
Edges versus Bends: Gravitational Accelerations in Filaments

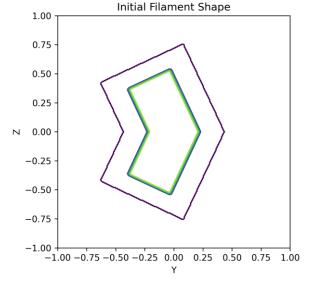
Rachel Curran

Department of Physics and Astronomy, University of North Carolina at Chapel Hill



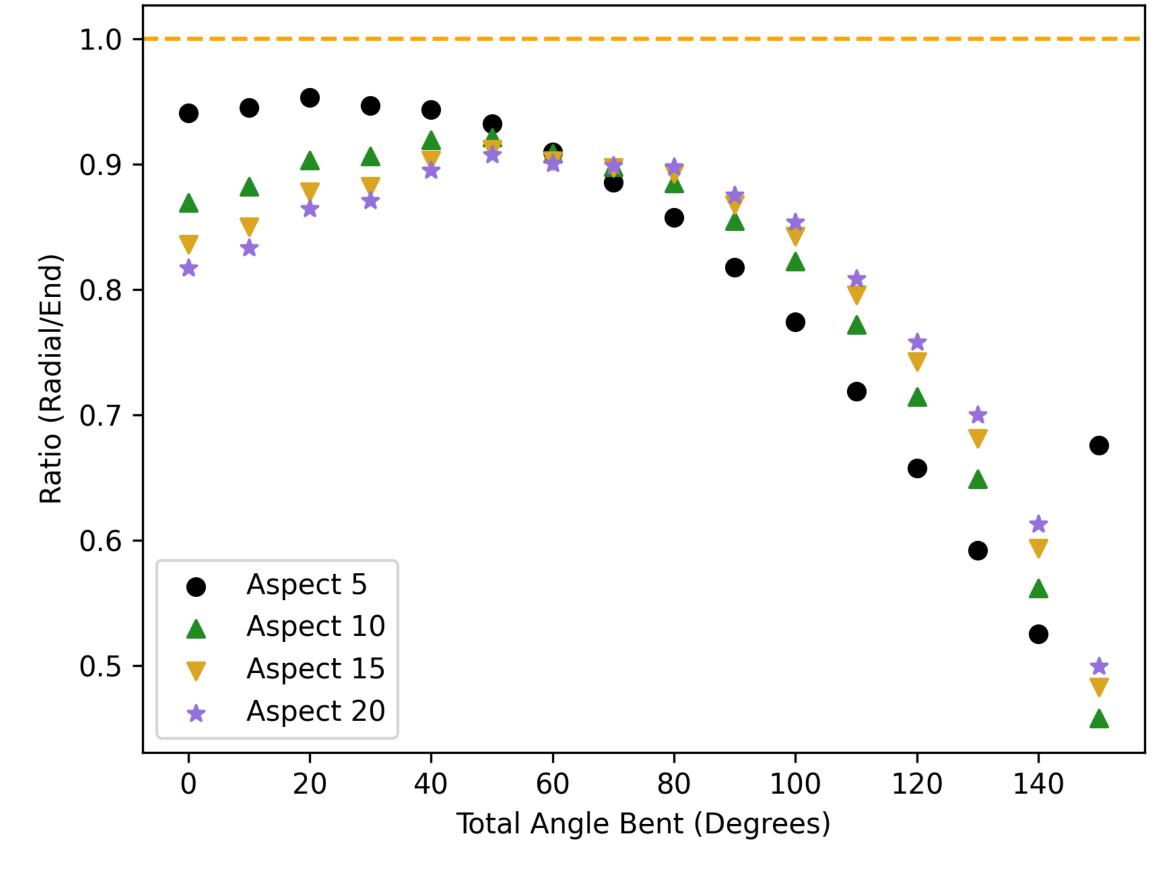
- > Tested different total bend angles at varying aspect ratios.
- > Aspect ratio is the ratio between the height and the radius.
- Total filament height is 1
- Used a uniform surface density





Can the accelerations at the bends "win" over the accelerations at the ends?

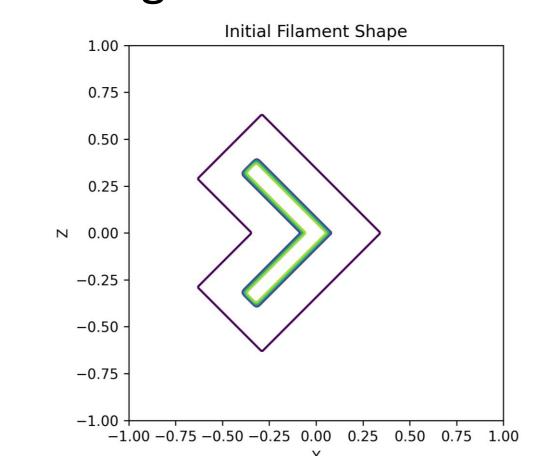
Ratio Between End and Center Accelerations



Conclusion and Future Work

- \succ Without smoothing out the ends, there are no angles where the bend causes a greater acceleration than the that at the ends.
- Repeat tests with realistic values for the initial densities instead of uniform values.
- > Find the accelerations for filaments with varying levels of smoothing of the ends.

Initial Conditions



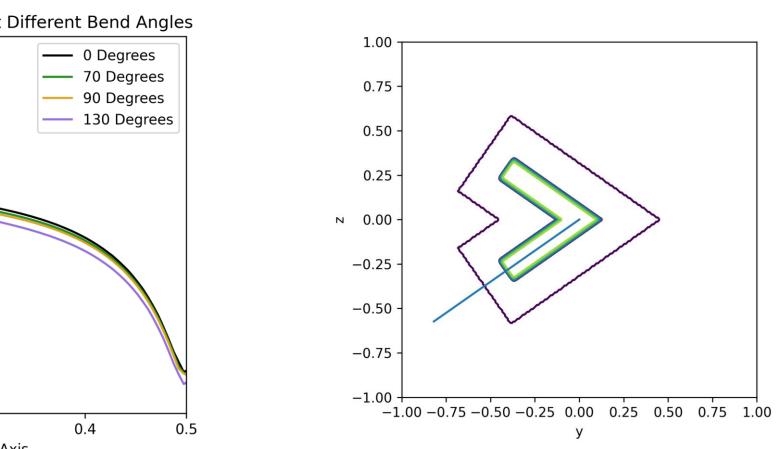
Accelerations Accelerations projected along one half of the bent filament, from the center to the end. > When there is no bend, the accelerations go to zero as they approach the bend. > However, with a bend in the filament, there is an opposing acceleration from the center that increases with the bend angle. — 0 Degrees 90 Degrees 130 Degree -0.05 -0.15-0.2Accelerations projected across the center of the filament. Again, with no bend this behaves as expected. > With some bend, the accelerations on the outer side of the filament increase in magnitude. However, as the bend increases and both ends approach each other, the mass from both halves of the filament contribute to the acceleration. Accelerations of Aspect 30 Filament at Different — 0 Degrees — 70 Degrees 90 Degrees — 130 Degrees -0.25 -0.10-0.15-0.20-1.00 -0.75 -0.50 -0.25 0.00 0.25 0.50 0.75 1.00

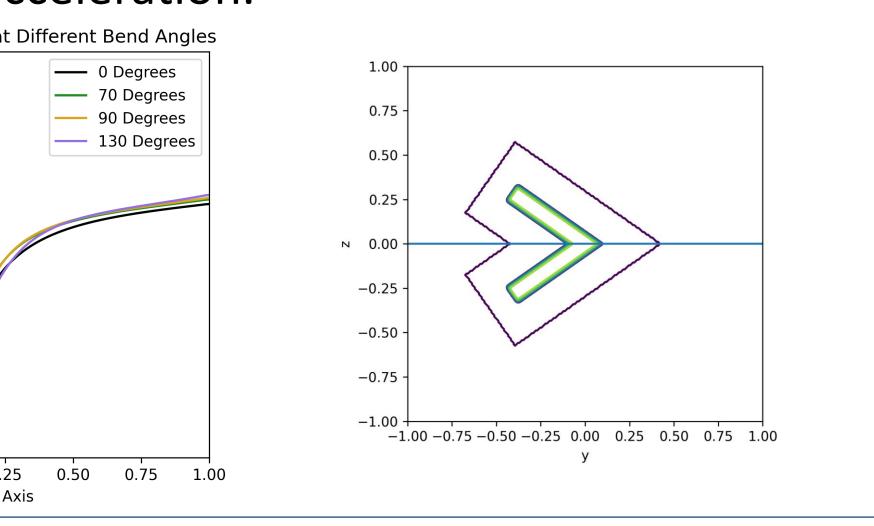
> -0.75 -0.50 -0.25 0.00 0.25 0.50 0.75 1.00 Length Along Center Axis

- Astrophysical Journal 616.1 (2004): 288.
- *Astrophysical Journal* 756.2 (2012): 145.
- Davide De Martin.

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I would like to thank my research advisor, Dr. Fabian Heitsch, as well as the members of the Astrophysical Fluid Dynamics Lab at UNC.





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2. Hoemann, Elena, Stefan Heigl, and Andreas Burkert. "Filament collapse: a two phase process." Monthly Notices of the Royal Astronomical Society 521.4 (2023): 5152-5159.

3. Pon, Andy, et al. "Aspect ratio dependence of the free-fall time for non-spherical symmetries." *The*

4. Image: ESO/APEX (MPIfR/ESO/OSO)/A. Hacar et al./Digitized Sky Survey 2. Acknowledgment:

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